

Muon and FFAG

Akira Sato
Osaka University

2015/09/17
FFAG2015 @ Kyusyu

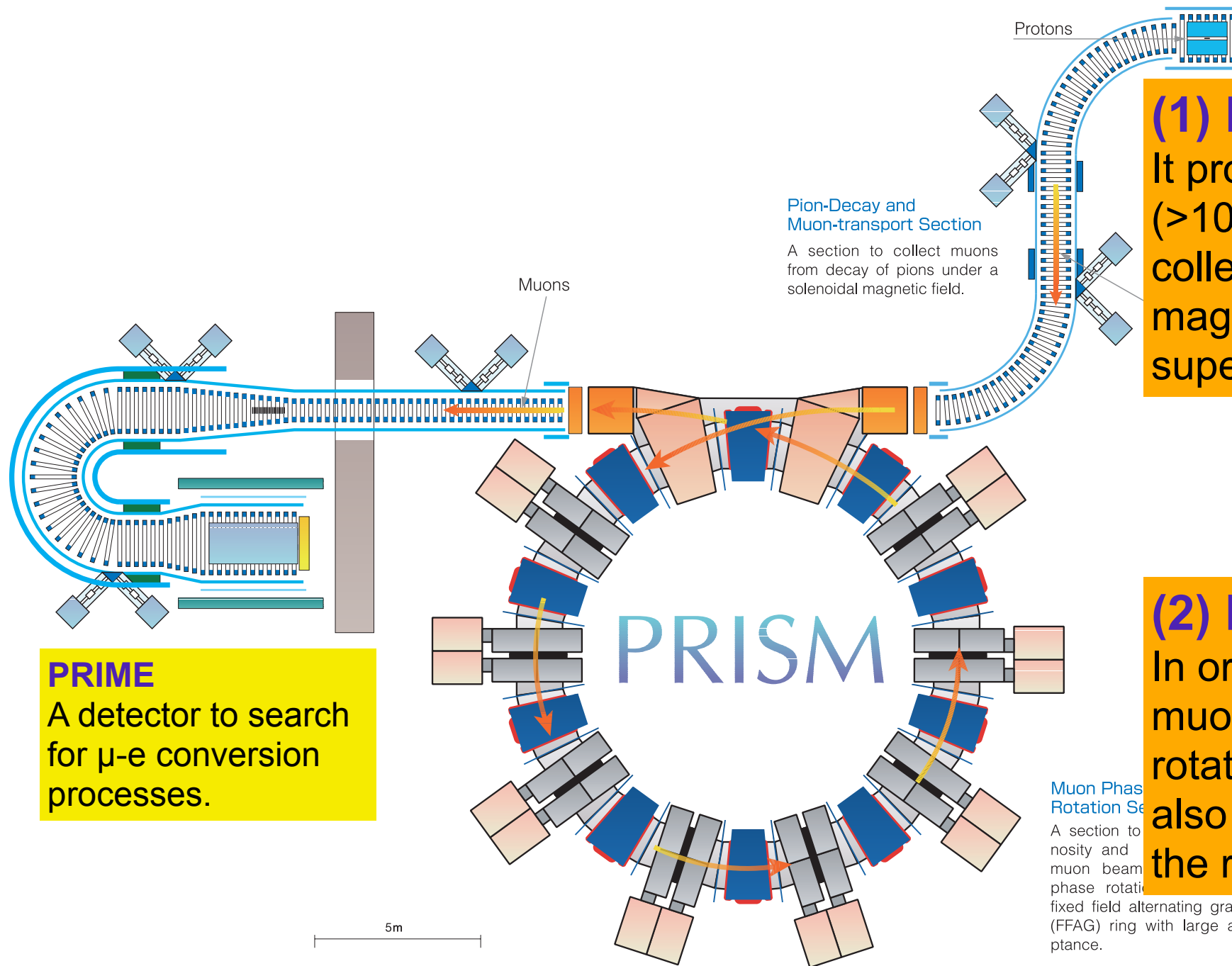
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- ◆ Status of my recent projects
 - MuSIC
 - COMET
 - PRISM/PRIME
- ◆ Ideas for μ -e conv. and nuSTORM/NF
- ◆ Some comments to FFAGs
 - Reliability of FFAG design performance
 - Perfect scaling FFAG

Status of my recent projects

- ♦ PRISM project is the origin of my interest in FFAG. It is an experiment to search the muon to electron conversion, which is strongly forbidden in the Standard Model of the particle physics, with the ultimate sensitivity.
- ♦ The experiment needs an very intense and monochromatic muon beam.
 - μ intensity : $10^{11} \sim 10^{12}$ μ -/sec
 - momentum : $\sim 40 \text{ MeV}/c$ with $\Delta p/p$ of $\sim \pm 2\%$
 - no pion contaminations

Key components of PRISM/PRIME



(1) Pion Capture System

It provides an intense muon beam ($>10^{12}\mu$ -/sec). Pions and muons are collected effectively by high solenoidal magnetic field produced by superconducting magnets.

(2) Muon Phase Rotator

In order to realize a monochromatic muon beam of $\Delta p/p \sim \pm 2\%$, a bunch rotation is applied in a FFAG ring. It also works to decay out the pions in the ring.

$$B(\mu^- + Ti \rightarrow e^- + Ti) < 10^{-18}$$

The 1st pion capture system : MuSIC

at RCNP, Osaka Univ. (Built 2009JFY)

Pion capture solenoid
Max. B_{sol}: 3.5 T

Pion-Muon transport solenoid (36deg.)
Max. B_{sol}: 2.0 T
Max. B_{dipole}: 0.04 T

Muons

WSS proton beam line
392MeV, 1μA

Successfully demonstrated to produce $10^8 \mu/\text{sec}$ muon beam with 400W (392MeV, 1μs) proton beam.

2 Aug. 2010

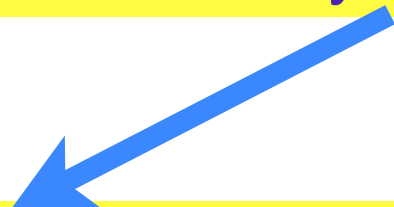
MuSIC: Development of the pion capture system

(1) Pion Capture System

It provides an intense muon beam ($>10^{12}\mu\text{-/sec}$). Pions and muons are collected effectively by high solenoidal magnetic field produced by superconducting magnets.



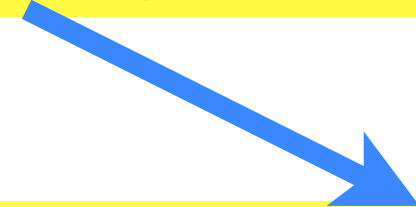
The 1st Pion Capture System, MuSIC, was developed and its performance was successfully demonstrated at RCNP, Osaka.



Start **COMET** experiment at J-PARC

A μ -e conversion experiment was funded by KEK/J-PARC. The experiment named COMET aims the sensitivity of 10^{-16} without the FFAG ring.

Working as
the main detector (CyDet) group leader



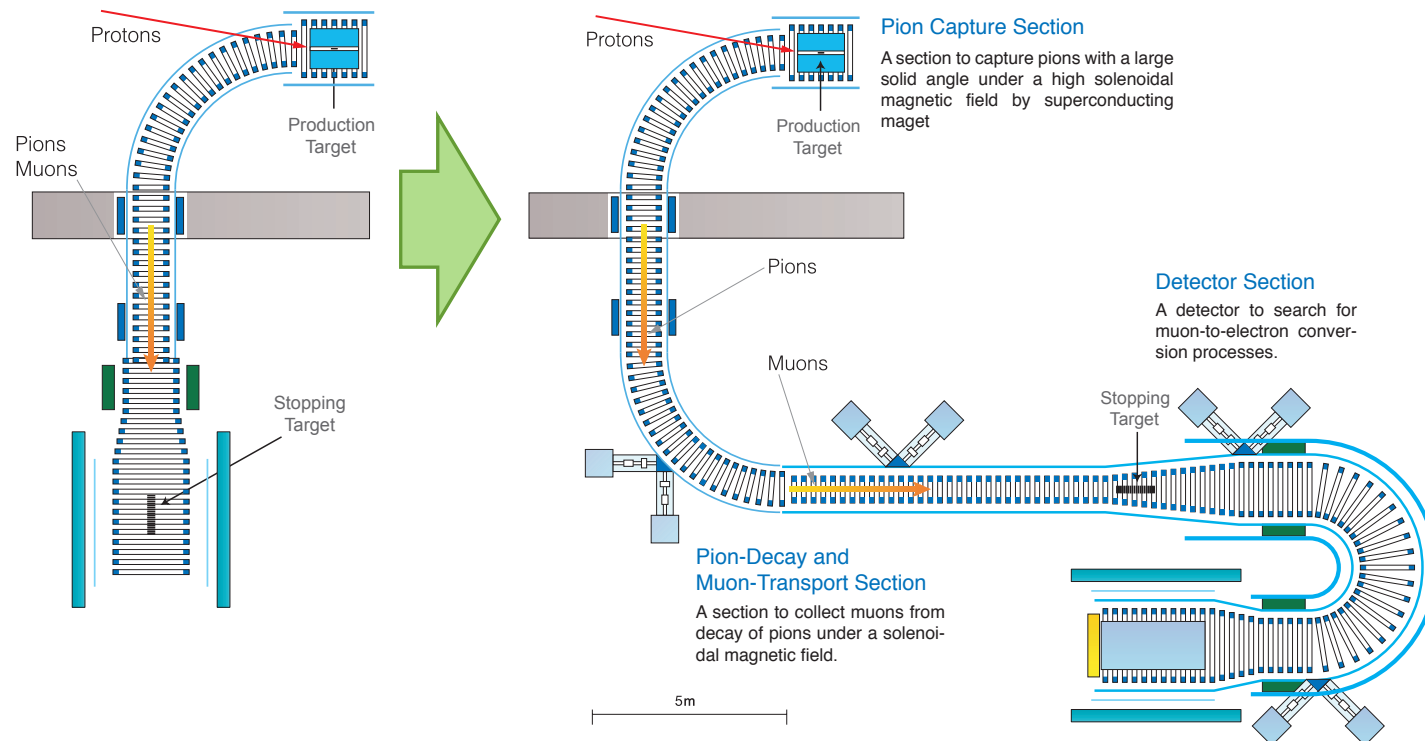
Start construction of μ beamline **MuSIC-M1** at RCNP

To start muon experiments for material science, nuclear physics chemical etc, the MuSIC solenoid channel was extended with a normal conducting beam line.

Working as
the responsible person of MuSIC

Staging approach of the μ -e conv. experiments

COMET @J-PARC



COMET Phase-I :

physics run 2017-

$BR(\mu+Al \rightarrow e+Al) < 7 \times 10^{-15}$ @ 90%CL

*8GeV-3.2kW proton beam, 12 days

*90deg. bend solenoid, cylindrical detector

*Background study for the phase2

COMET Phase-II :

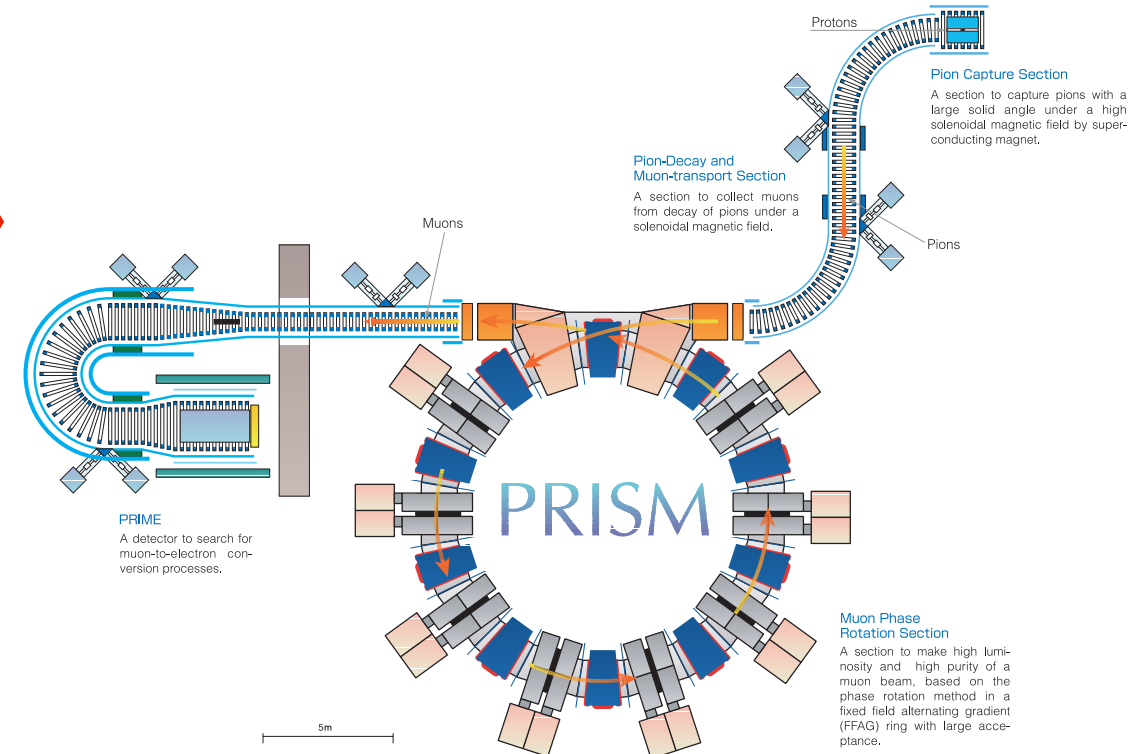
physics run 2020-

$BR(\mu+Al \rightarrow e+Al) < 6 \times 10^{-17}$ @ 90%CL

*8GeV-56kW proton beam, 2 years

*180deg. bend solenoid, bend spectrometer,
transverse tracker+calorimeter

PRISM



PRISM/PRIME :

physics run 202X-

$BR(\mu+Ti \rightarrow e+Ti) < 10^{-18}$ @ 90%CL

*4MW proton beam, 2 years

*180deg. bend solenoid, bend spectrometer,
transverse tracker+calorimeter

* FFAG as a phase rotator

Funded!

COMET : Collaboration

164 collaborators, 37 institutes



The COMET Collaboration (Sep. 2014)

R. Akhmetshin^{6,28}, V. Anishchik⁴, M. Aoki²⁹, R. B. Appleby^{8,22}, Y. Arimoto¹⁵, Y. Bagaturia³³, Y. Ban³, W. Bertsche²², A. Bondar^{6,28}, S. Canfer³⁰, S. Chen²⁵, Y. E. Cheung²⁵, B. Chiladze³², D. Clarke³⁰, M. Danilov^{13,23}, P. D. Dauncey¹¹, J. David²⁰, W. Da Silva²⁰, C. Densham³⁰, G. Devidze³², P. Dornan¹¹, A. Drutskey^{13,23}, V. Duginov¹⁴, A. Edmonds³⁵, L. Epshteyn^{6,27}, P. Evtoukhovich¹⁴, G. Fedotov^{6,28}, M. Finger⁷, M. Finger Jr⁷, Y. Fujii², Y. Fukao¹⁵, J-F. Genat²⁰, M. Gersabeck²², E. Gillies¹¹, D. Grigoriev^{6,27,28}, K. Gritsay¹⁴, E. Hamada¹⁵, R. Han¹, K. Hasegawa¹⁵, I. H. Hasim²⁹, O. Hayashi²⁹, M. I. Hossain¹⁶, Z. A. Ibrahim²¹, Y. Igarashi¹⁵, F. Ignatov^{6,28}, M. Iio¹⁵, M. Ikeno¹⁵, K. Ishibashi¹⁹, S. Ishimoto¹⁵, T. Itahashi²⁹, S. Ito²⁹, T. Iwami²⁹, Y. Iwashita¹⁷, X. S. Jiang², P. Jonsson¹¹, V. Kalinnikov¹⁴, F. Kapusta²⁰, H. Katayama²⁹, K. Kawagoe¹⁹, V. Kazanin^{6,28}, B. Khazin^{6,28}, A. Khvedelidze¹⁴, M. Koike³⁶, G. A. Kozlov¹⁴, B. Krikler¹¹, A. Kulikov¹⁴, E. Kulish¹⁴, Y. Kuno²⁹, Y. Kuriyama¹⁸, Y. Kurochkin⁵, A. Kurup¹¹, B. Lagrange^{11,18}, M. Lancaster³⁵, H. B. Li², W. G. Li², A. Liparteliani³², R. P. Litchfield³⁵, P. Loveridge³⁰, G. Macharashvili¹⁴, Y. Makida¹⁵, Y. Mao³, O. Markin¹³, Y. Matsumoto²⁹, T. Mibe¹⁵, S. Mihara¹⁵, F. Mohamad Idris²¹, K. A. Mohamed Kamal Azmi²¹, A. Moiseenko¹⁴, Y. Mori¹⁸, N. Mosulishvili³², E. Motuk³⁵, Y. Nakai¹⁹, T. Nakamoto¹⁵, Y. Nakazawa²⁹, J. Nash¹¹, M. Nioradze³², H. Nishiguchi¹⁵, T. Numao³⁴, J. O'Dell³⁰, T. Ogitsu¹⁵, K. Oishi¹⁹, K. Okamoto²⁹, C. Omori¹⁵, T. Ota³¹, H. Owen²², C. Parkes²², J. Pasternak¹¹, C. Plostinar³⁰, V. Ponariadov⁴, A. Popov^{6,28}, V. Rusinov^{13,23}, A. Ryzhenenkov^{6,28}, B. Sabirov¹⁴, N. Saito¹⁵, H. Sakamoto²⁹, P. Sarin¹⁰, K. Sasaki¹⁵, A. Sato²⁹, J. Sato³¹, D. Shemyakin^{6,28}, N. Shigyo¹⁹, D. Shoukavy⁵, M. Slunicka⁷, M. Sugano¹⁵, Y. Takubo¹⁵, M. Tanaka¹⁵, C. V. Tao²⁶, E. Tarkovsky^{13,23}, Y. Tevzadze³², N. D. Thong²⁹, V. Thuan¹², J. Tojo¹⁹, M. Tomasek⁹, M. Tomizawa¹⁵, N. H. Tran²⁹, I. Trek³², N. M. Truong²⁹, Z. Tsamalaidze¹⁴, N. Tsverava¹⁴, S. Tygier²², T. Uchida¹⁵, Y. Uchida¹¹, K. Ueno¹⁵, S. Umasankar¹⁰, E. Velicheva¹⁴, A. Volkov¹⁴, V. Vrba⁹, W. A. T. Wan Abdullah²¹, M. Warren³⁵, M. Wing³⁵, T. S. Wong²⁹, C. Wu^{2,25}, G. Xia²², H. Yamaguchi¹⁹, A. Yamamoto¹⁵, M. Yamanaka²⁴, Y. Yang¹⁹, H. Yoshida²⁹, M. Yoshida¹⁵, Y. Yoshii¹⁵, T. Yoshioka¹⁹, Y. Yuan², Y. Yudin^{6,28}, J. Zhang², Y. Zhang²

¹North China Electric Power University, Beijing, People's Republic of China

²Institute of High Energy Physics (IHEP), Beijing, People's Republic of China

³Peking University, Beijing, People's Republic of China

⁴Belarusian State University (BSU), Minsk, Belarus

⁵B.I. Stepanov Institute of Physics, National Academy of Sciences of Belarus, Minsk, Belarus

⁶Budker Institute of Nuclear Physics (BINP), Novosibirsk, Russia

⁷Charles University, Prague, Czech Republic

⁸The Cockcroft Institute, Daresbury Laboratory, Warrington, UK

⁹Czech Technical University, Prague, Czech Republic

¹⁰Indian Institute of Technology, Bombay, India

¹¹Imperial College London, London, UK

¹²Institute for Nuclear Science and Technology, Hanoi, Vietnam

¹³Institute for Theoretical and Experimental Physics (ITEP), Russia

¹⁴Joint Institute for Nuclear Research (JINR), Dubna, Russia

¹⁵High Energy Accelerator Research Organization (KEK), Tsukuba, Japan

¹⁶King Abdulaziz University, Saudi Arabia

¹⁷Institute for Chemical Research, Kyoto University, Kyoto, Japan

¹⁸Research Reactor Institute, Kyoto University, Kyoto, Japan

¹⁹Kyushu University, Fukuoka, Japan

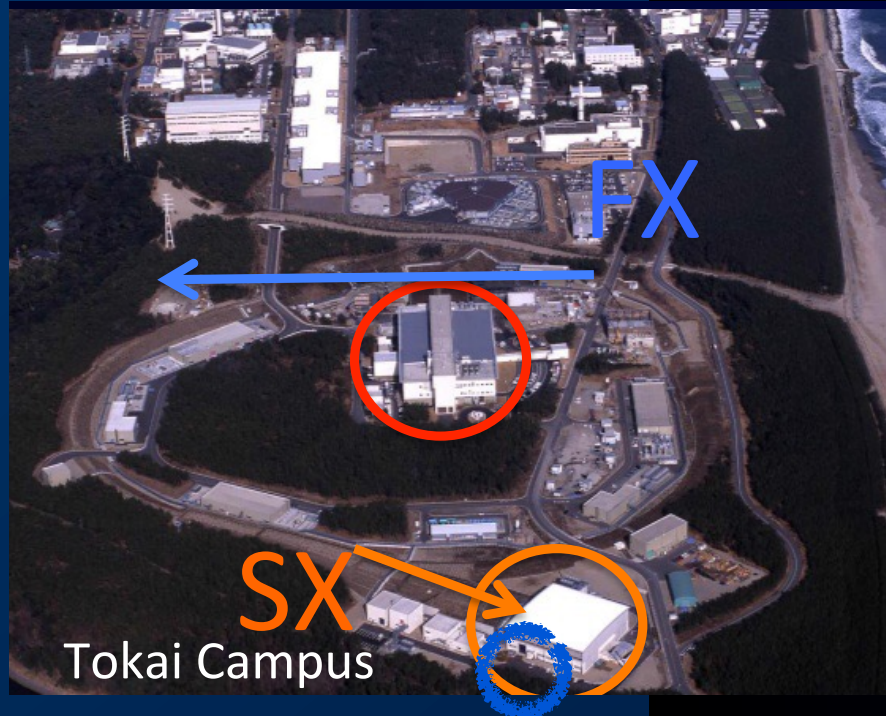
²⁰Laboratory of Nuclear and High Energy Physics (LPNHE), CNRS-IN2P3 and University Pierre and Marie Curie (UPMC), Paris, France

²¹National Centre for Particle Physics, Universiti Malaya, Kuala Lumpur, Malaysia

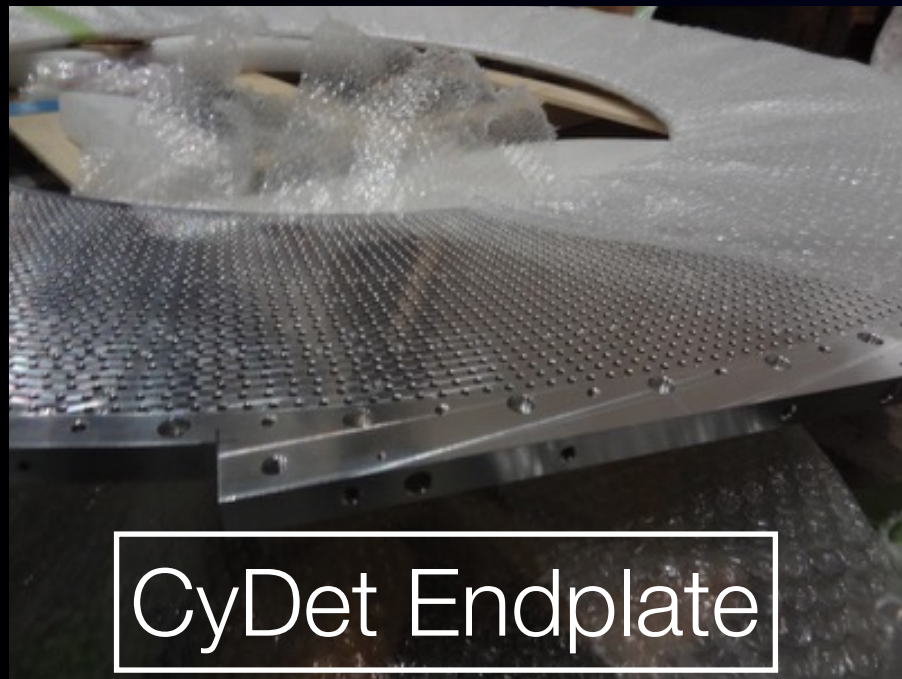
Me

COMET : Status

- COMET building completed!



COMET : Status | CyDet



Cylindrical Drift Chamber for e^- tracking

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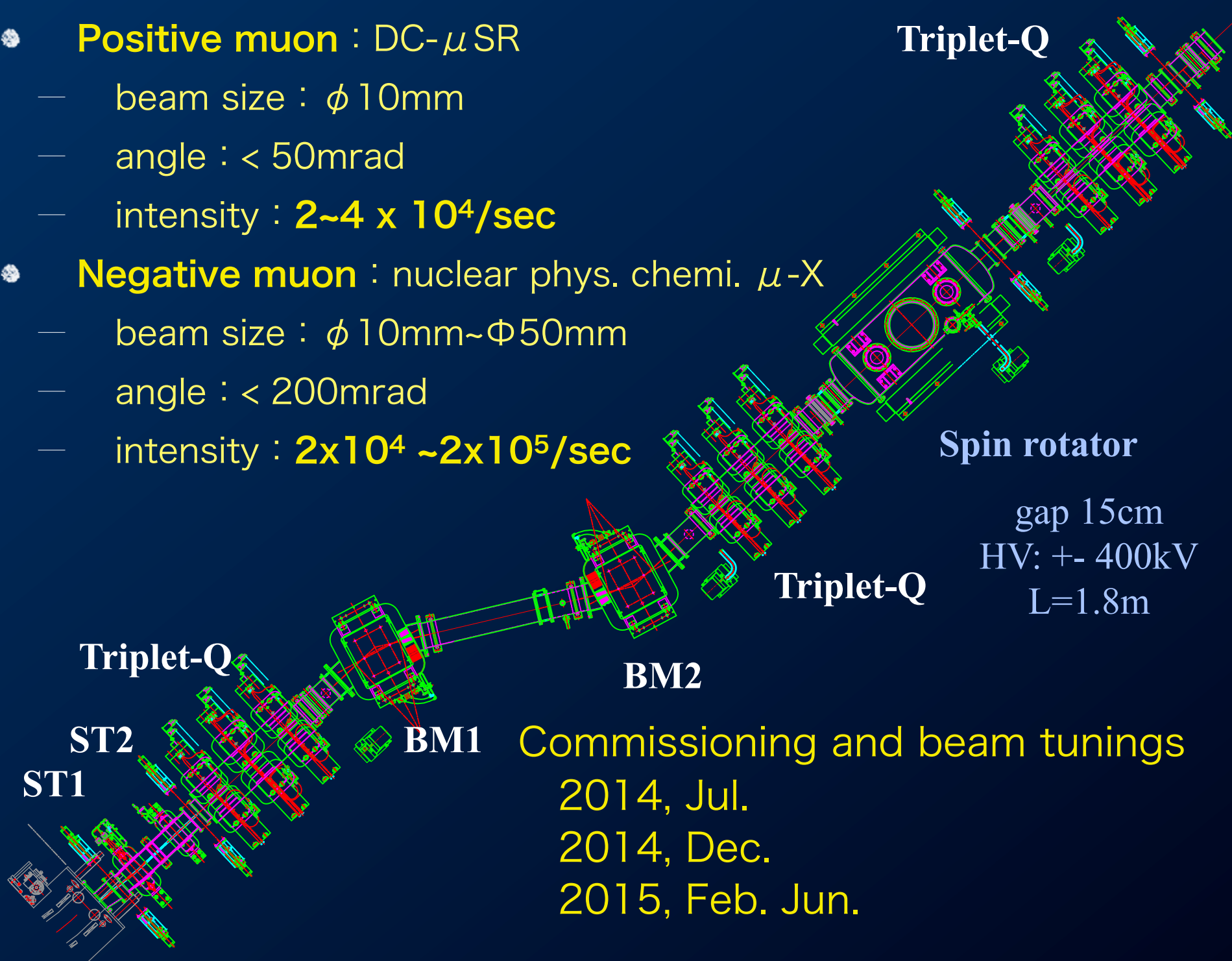
Working as
the responsible person of MuSIC

The DC muon beam line in Japan

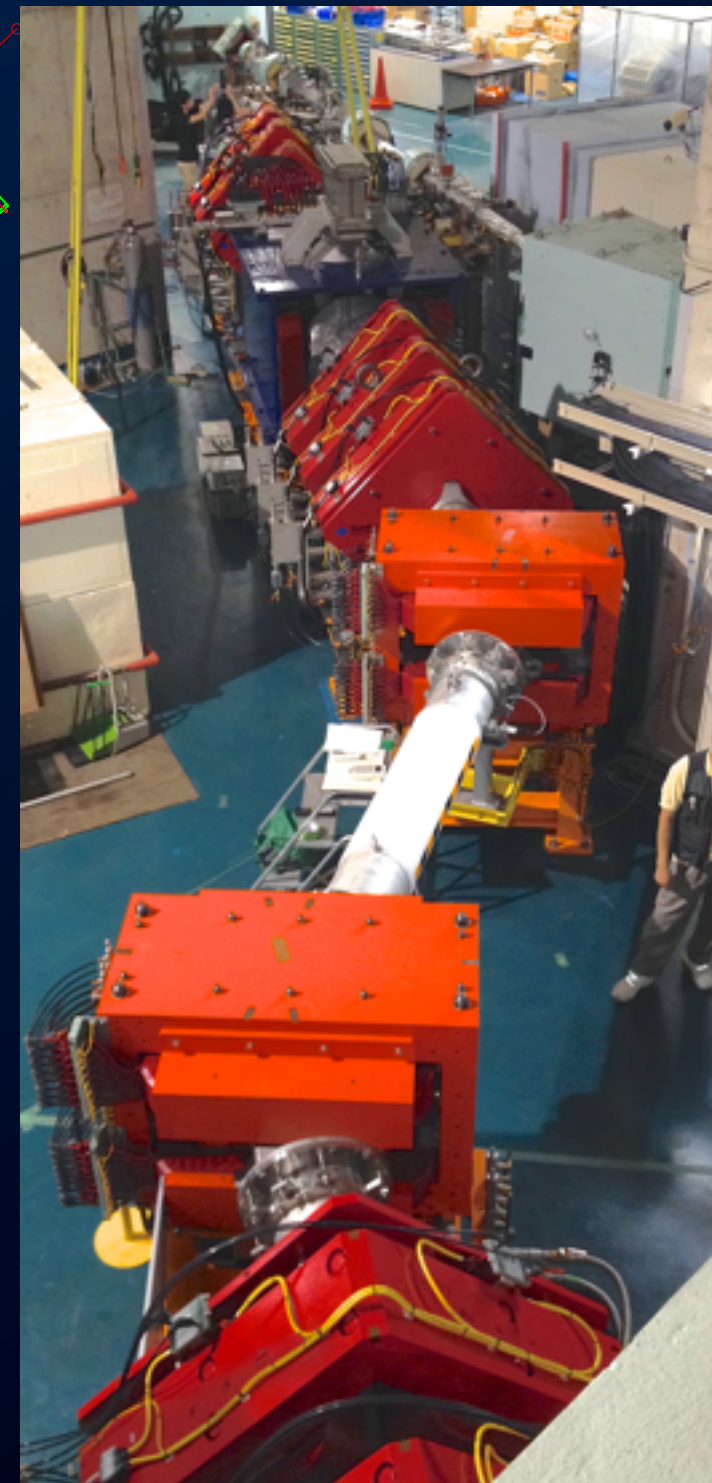
RCNP-MuSIC-M1 build in 2013JFY

Goal of the beam performance

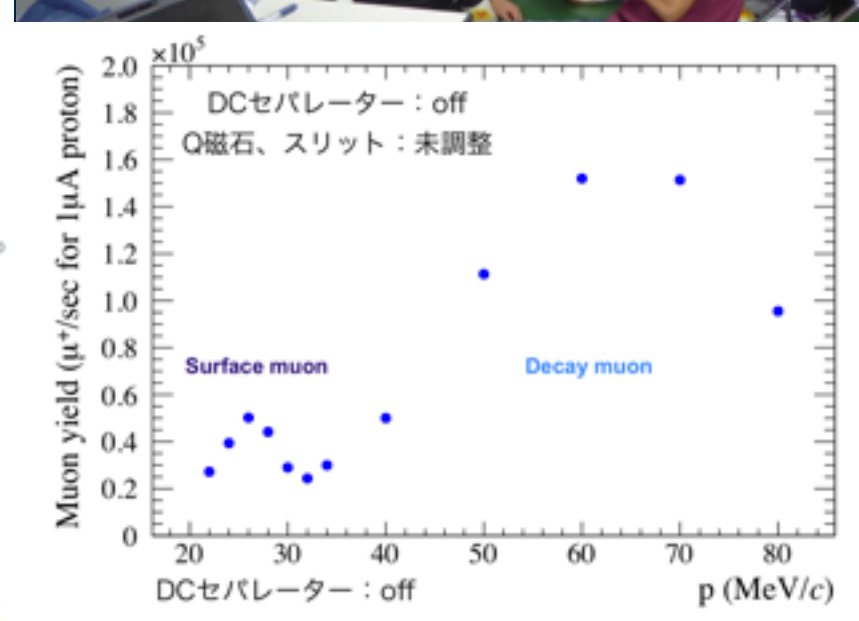
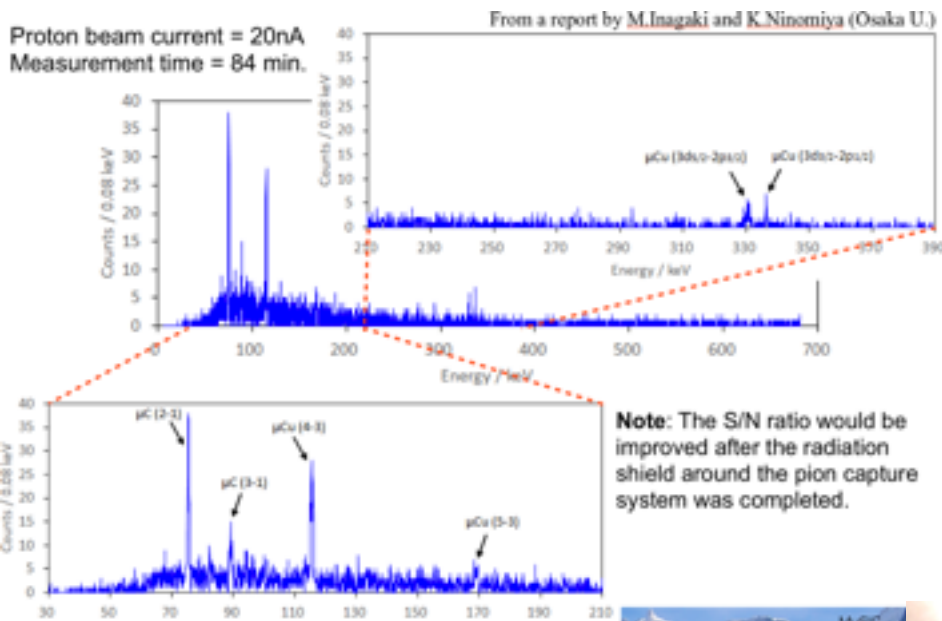
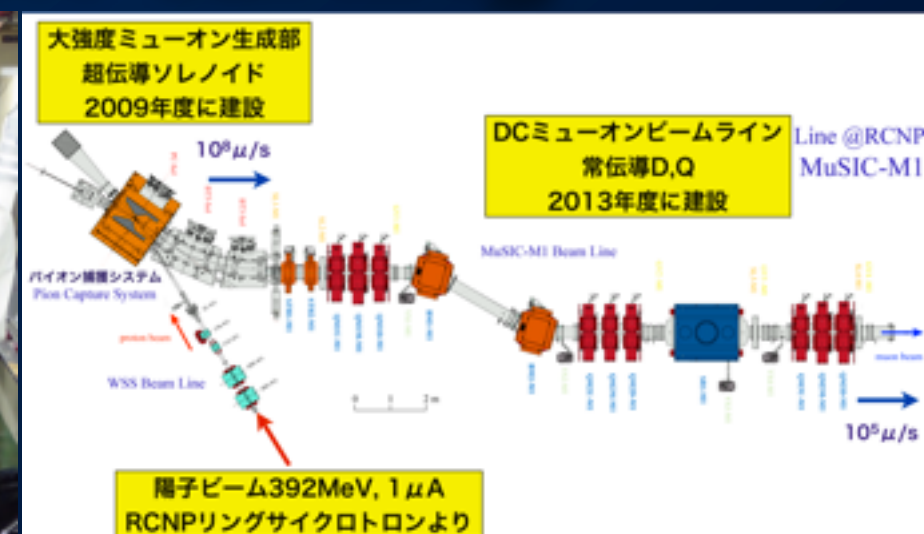
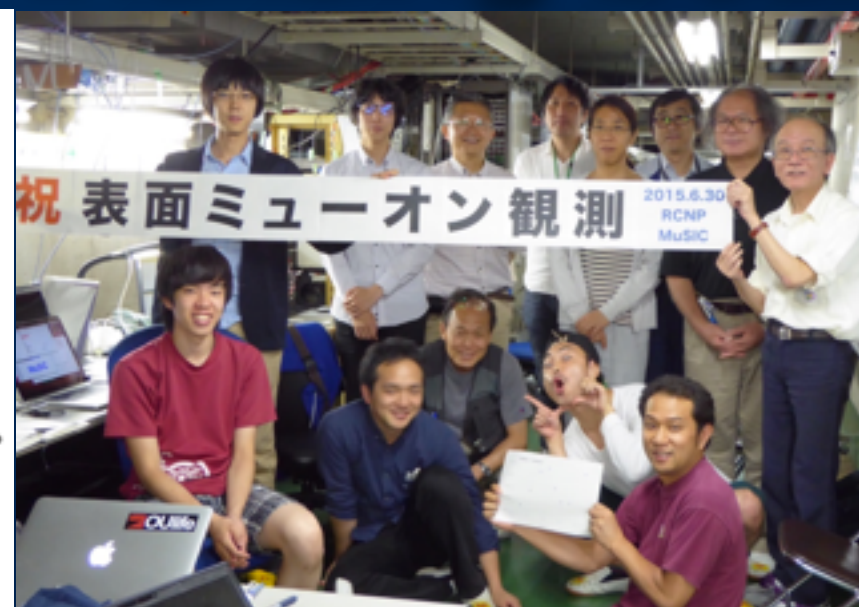
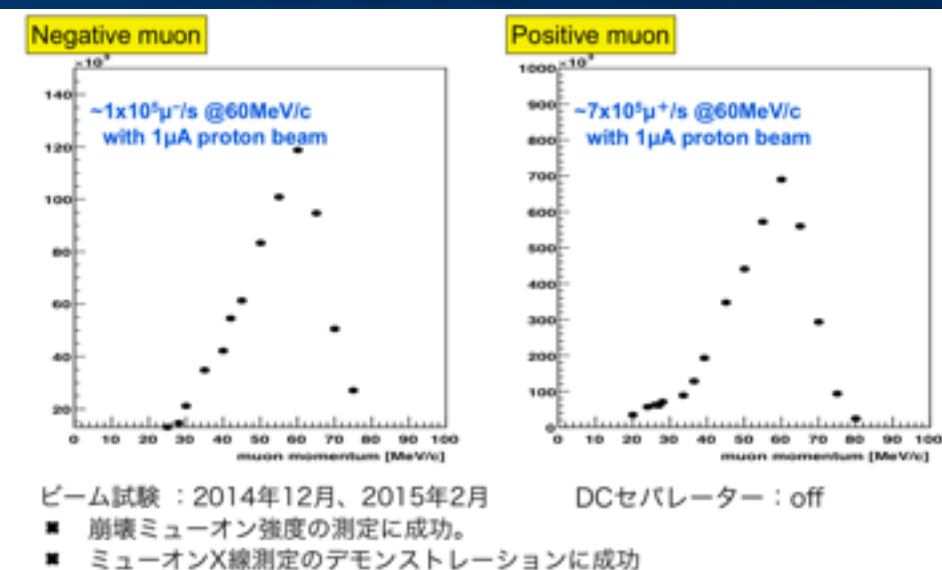
- **Positive muon** : DC- μ SR
 - beam size : ϕ 10mm
 - angle : $< 50\text{mrad}$
 - intensity : $2\sim 4 \times 10^4/\text{sec}$
- **Negative muon** : nuclear phys. chemi. μ -X
 - beam size : ϕ 10mm~ ϕ 50mm
 - angle : $< 200\text{mrad}$
 - intensity : $2 \times 10^4 \sim 2 \times 10^5/\text{sec}$



Commissioning and beam tunings
2014, Jul.
2014, Dec.
2015, Feb. Jun.

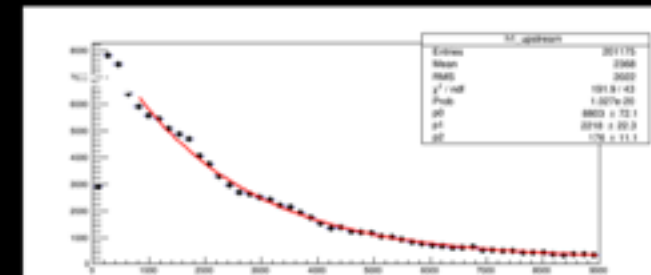


MuSIC-M1: Finalizing beam tuning

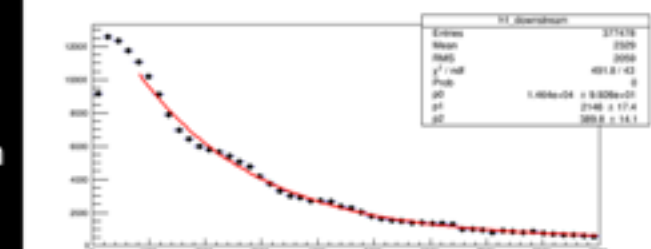


崩壊ミューオンビーム (60 MeV/c) による

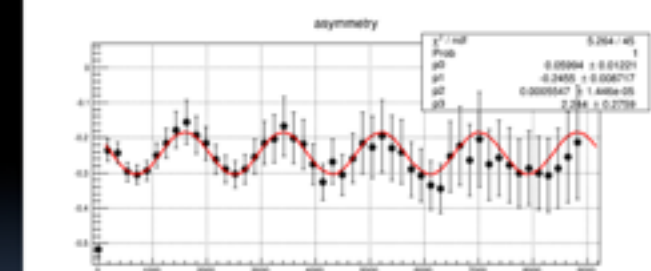
upstream
counter :U



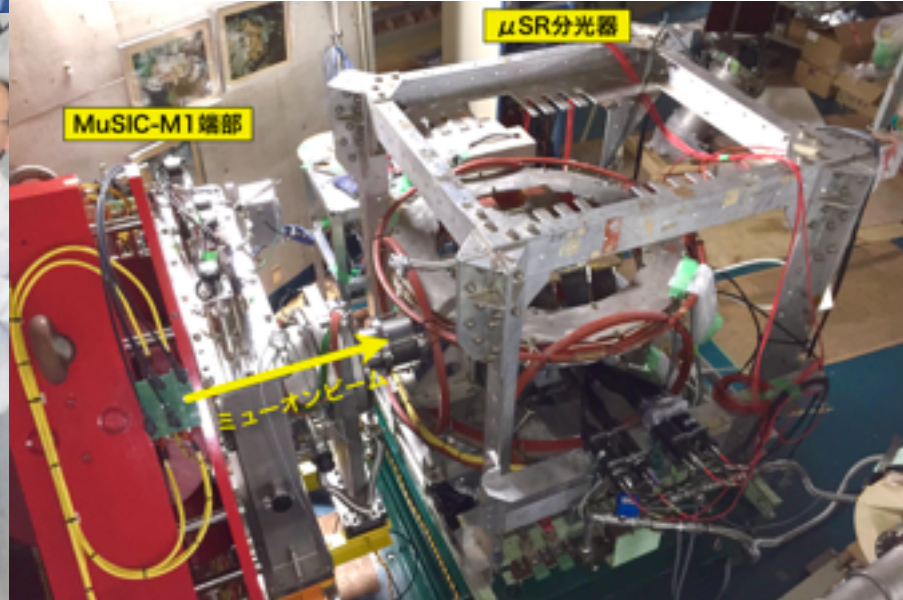
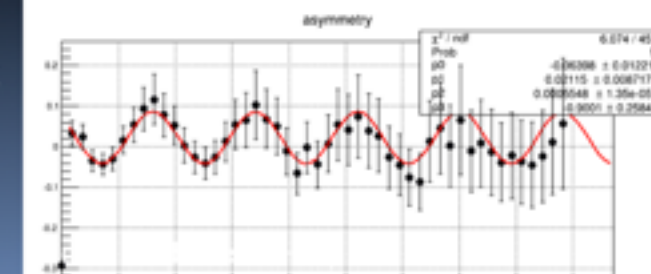
downstream
counter :D



Asymmetry



ベースラインがゼロとなるようα補正を入れる

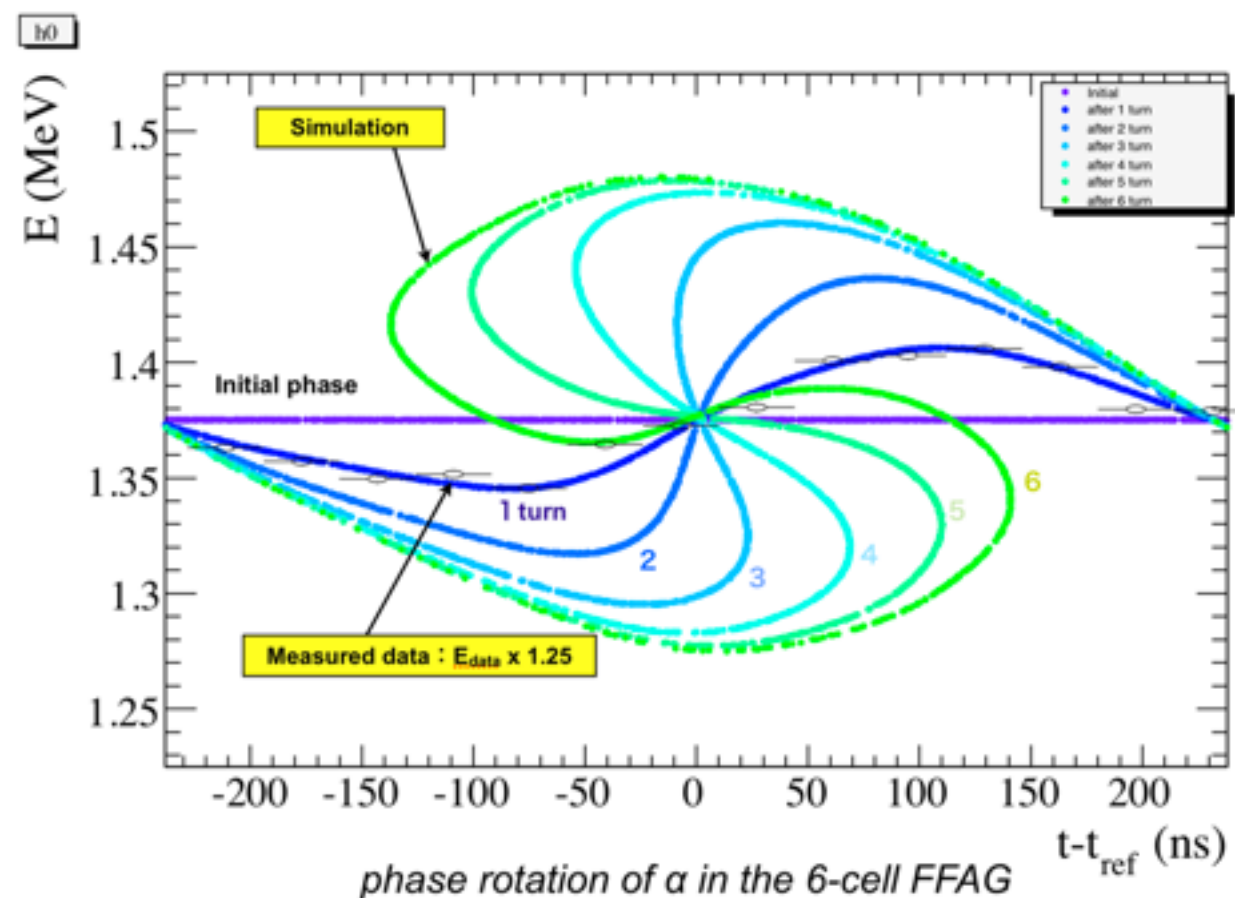


Toward the PRISM

- ultimate μ -e
experiment

PRISM-FFAG R&D

R&Ds@RCNP
2003.4-2009.3

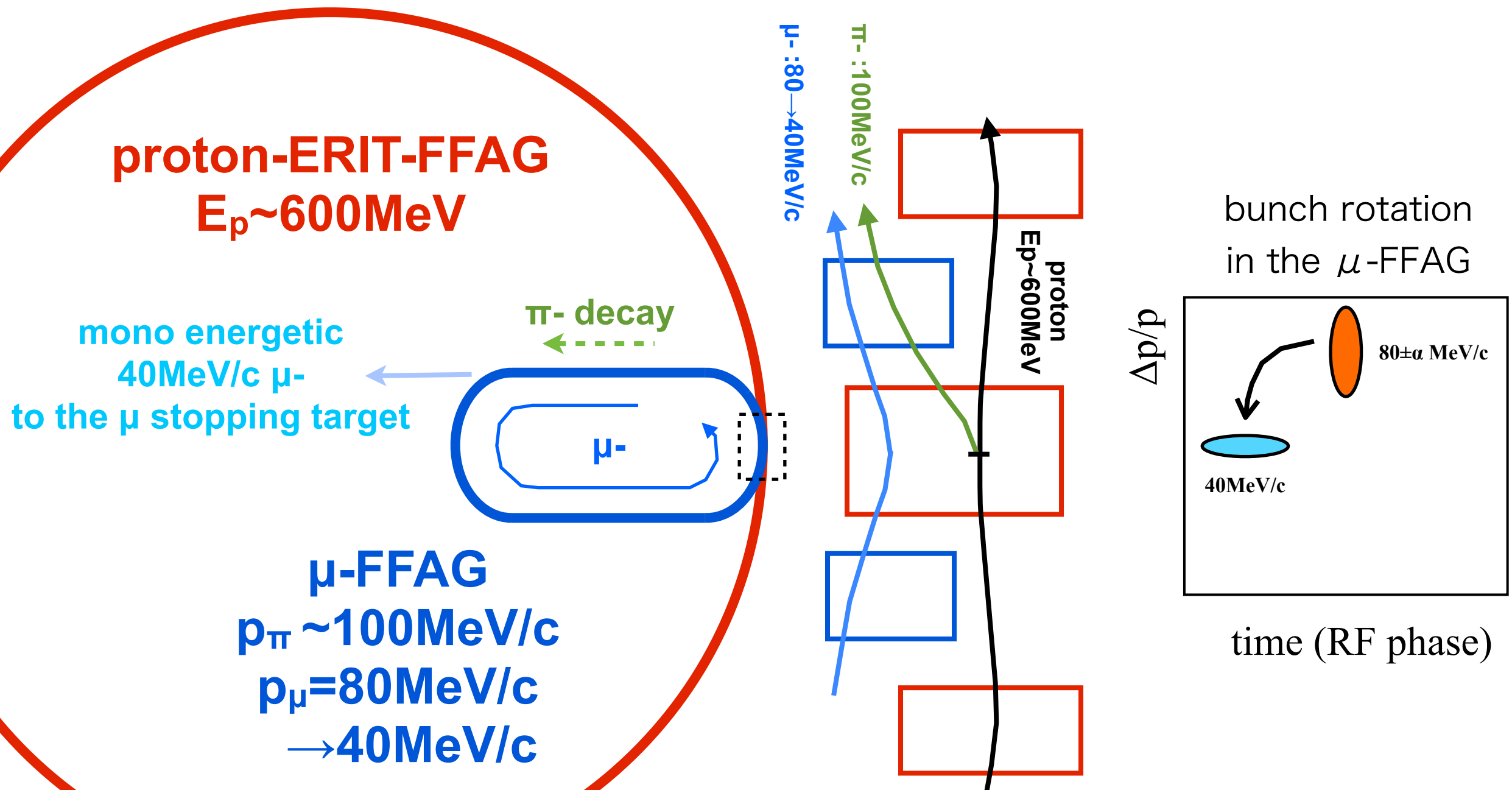


- PRISM-FFAG was disassembled. It is sleeping at W-exp hall of RCNP.
- The PRISM task force was formed (Jaroslow's talk).

Issues of the PRISM

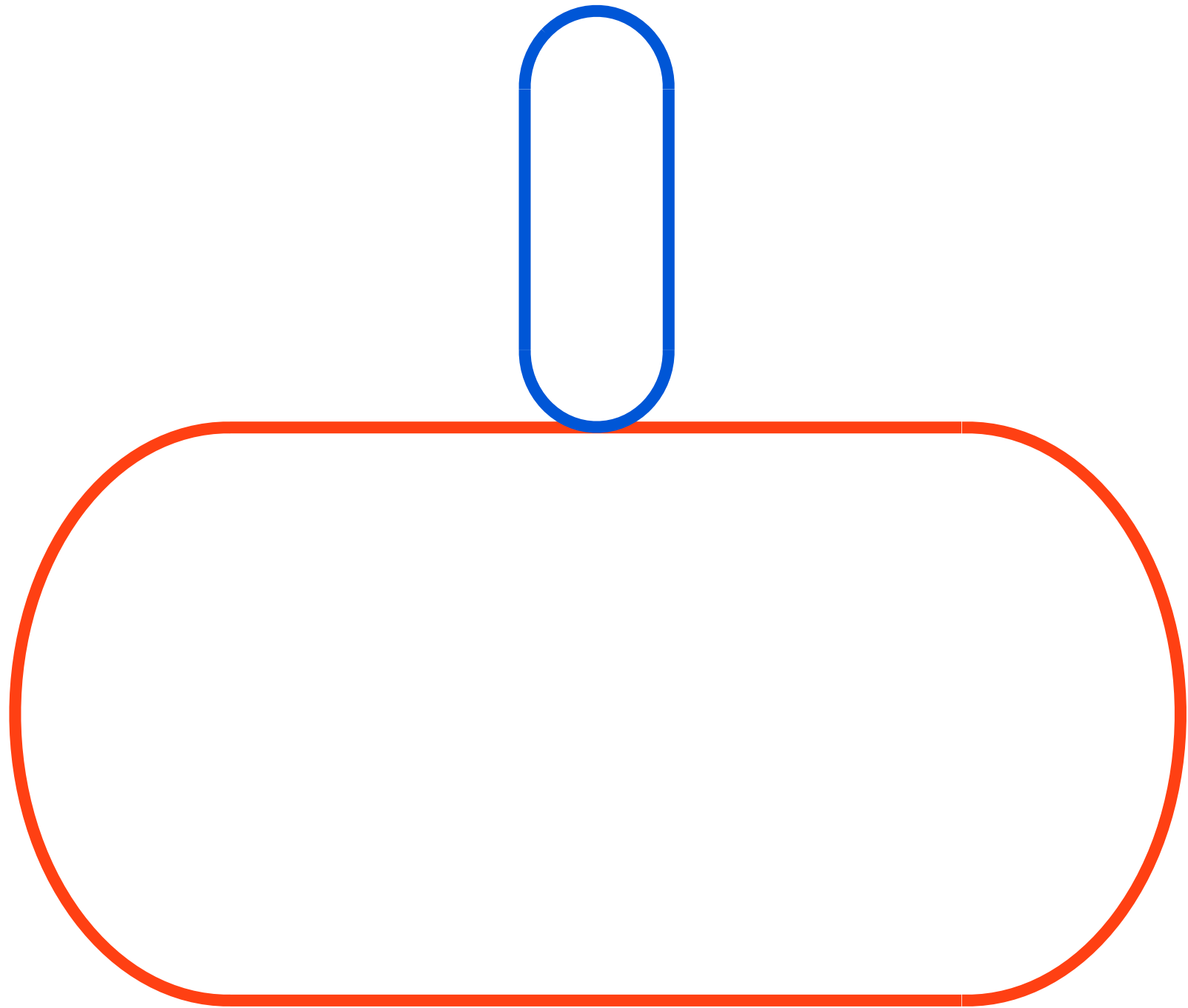
- ◆ There are mainly two issues:
 - the transport line from the solenoidal pion decay channel into the FFAG ring
 - the injection and extraction system.
- ◆ Need to solve
- ◆ I will show conceptual figures to solve some of these issues.

A concept of new PRISM



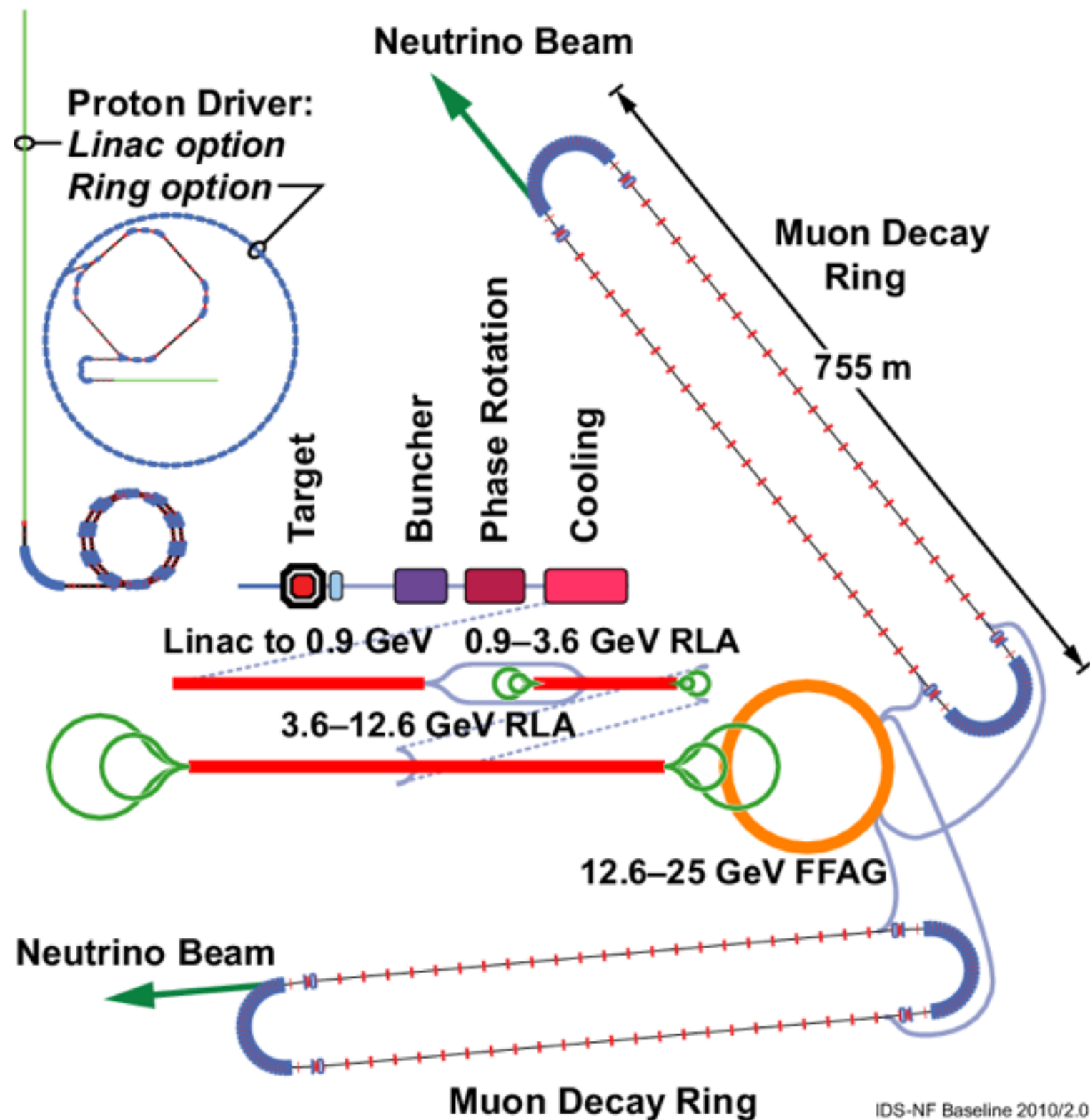
- ✦ Muons are decelerated in the μ^- -FFAG ring.
- ✦ V-FFAG or H-FFAG?
- ✦ Good solutions for extraction?

A concept of new PRISM



- ◆ A racetrack FFAG is better for the proton ring?

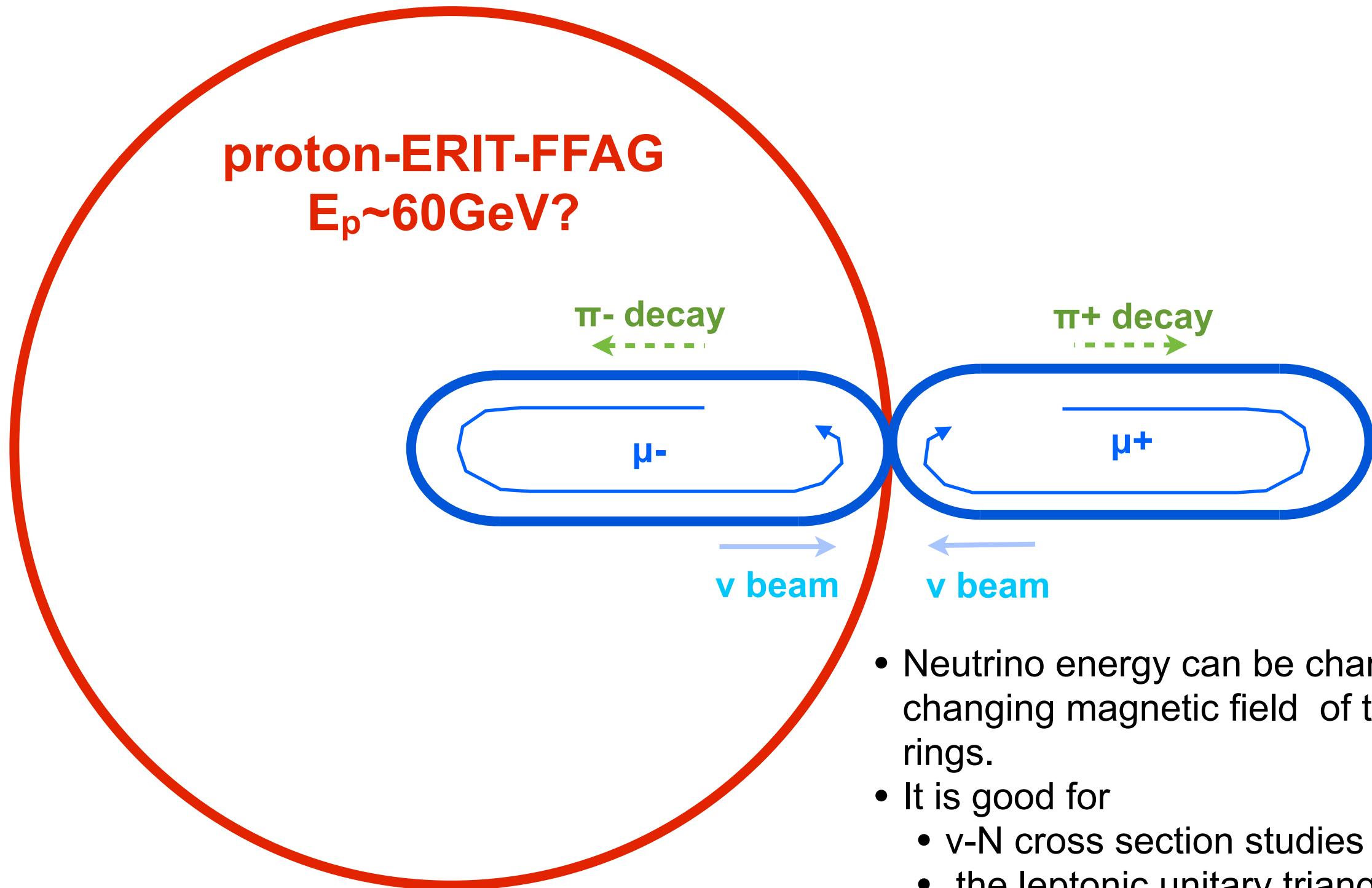
IDS-NF baseline 2010/2.0



- ◆ After the pion production target, it has
 - buncher,
 - phase rotator,
 - cooling, and
 - muon accelerator ...
- ◆ Very complicated and a lot of technical challenges !

IDS-NF Baseline 2010/2.0

MERIT + nuSTORM / NF



- Neutrino energy can be changed by changing magnetic field of the μ -FFAG rings.
- It is good for
 - ν -N cross section studies and
 - the leptonic unitary triangle studies.

✦ possible?

✦ Is this work as a NF? μ acceleration is not necessary!

Some comments to FFAGs

Comment 1

- ◆ We have built many FFAGs and studied many things. We also have many unique ideas which work with FFAG.
- ◆ I think it is very important time for future FFAGs, because some projects are considering to use FFAGs as practical accelerators. eRICH, PRISM,,
- ◆ Some people said...
 - “We can not reproduce performance of real built FFAG in computer models”, pointed out by Meot.
 - It is very crucial problem. Performance of practical accelerators can be predicted by simulations.
 - “We need to demonstrate many aspects of FFAGs”

Study with booster or ERIT

- ♦ Most of people are working with 150MeV-FFAG at KURRI and Kyushyu-U.
 - There are many progress in these years.
 - But, I think still we have a lot of unknown things of the 150-MeV FFAGs.
 - It is the most difficult FFAG,
 - yoke free, saturation in the iron
 - 500Gauss stray field at the center of straight section.
 - There are many 150MeV-FFAG specific issues.
 - We should categorize the issues to
 - (1) FFAG common issues
 - solve and study the issues with the booster FFAG / ERIT FFAG
 - (2) 150MeV specific issues
 - need to work with 150MeV FFAG
 -

Scaling FFAGs operated

	Site	Ion	E _{max} (MeV)	Sector type	Cell	Radius (m)	First Beam	Acc. Scheme	Comments	
	PoP	KEK	p	1	Radial	8	0.8-1-1	2000	rf	The 1st proton FFAG. Accelerate protons from 50 keV to 500 keV in 1 msec.
	150MeV	KEK	p	150	Radial	12	4.5-5.2	2003	rf	100 Hz, 90% extraction
	ADSR-Injector	KURRI	p	2.5	Spiral (40°)	8	0.6-1.0	2006	induction	Initial spec : 120 Hz, 1 μA Later : 1 kHz, 100 μA, 200 MeV
	ADSR-Booster			20	Radial	8	1.4-1.7	2006	rf	
	ADSR-MR			150	Radial	12	4.5-5.1	2008	rf	
	ERIT	KURRI	p	11	Radial	8	2.35	2008	rf	70 mA ionization cooling ring to produce a neutron beam for BNCT
	PRISM-demo	Osaka-U.	α	1.0	Radial	6	3.0	2008	rf	Demonstrate phase-rotation for the mu-e conv. experiment (PRISM)
	EB-FFAG	NHV Co.	e	0.5	Spiral (30°)	6	0.19-0.44	2008	induction	Prototype of electron FFAG for industrial applications
	Radiatron	RadiaBeam Technologies, LLC (US)	e	5	Radial	12	0.3-0.7	(2008)	induction	Prototype of a high power accelerator for industrial applications

Scaling FFAG

- ◆ Features of the scaling-FFAG
 - beam dynamics dose not depends on the energy
 - zero-chromaticity (constant tune)
- ◆ Magnetic field : $B(r, \theta) = B_0 \left(\frac{r}{r_0} \right)^k f \left(\theta - \varsigma \ln \frac{r}{r_0} \right)$
 - usually we adjust B at the mid plane.
- ◆ But most of FFAG has tune variation
- ◆ I propose to make a magnetic field to realize a perfect scaling FFAG.